CONSEQUENCES OF RADIATION TRAPPING ON ELECTRON ENERGY DISTRIBUTIONS IN LOW PRESSURE INDUCTIVELY COUPLED Hg/Ar DISCHARGES*

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AGENDA

- Radiation transport in low pressure plasmas
- Overview of the Hybrid Plasma Equipment Model
- The Monte Carlo Radiation Transport Model
- Effect of radiation trapping on EEDFs
- Variation of EEDF with
 - Power
 - Frequency
 - Radiative lifetime
- Conclusions

ELECTRODELESS LAMPS AND TRAPPING

- Electrodeless gas discharges are finding increasing use in the lighting industry due to their increased lifetime.
- Investigations are underway to increase the efficiency of these lamps, now \cong 25%.
- Typical fluorescent lamps consist of Ar/Hg ≈ 97/3. Resonance radiation from Hg (6³P₁) (254 nm) and Hg (6¹P₁) (185 nm) excites phosphors which generate visible light.
- This resonance radiation can be absorbed and reemitted many times prior to striking the phosphor, increasing the effective lifetime of emission as viewed from outside the lamp.
- We have modeled this mechanism using a Monte Carlo radiation transport model linked to a hybrid plasma equipment model, to realistically simulate the discharge.



HYBRID PLASMA EQUIPMENT MODEL

- A modular simulator addressing low temperature, low pressure plasmas.
- EMM: electromagnetic fields and magneto-static fields
- EETM: electron temperature, electron impact sources, and transport coefficients
- FKM: densities, momenta, and temperatures of charged and neutral plasma species; and electrostatic potentials



REACTION CHEMISTRY

 $Rg + e \rightarrow Rg^* + e$ $Rg + e \rightarrow Rg^+ + 2e$ $Rg^* + e \rightarrow Rg^+ + 2e$ $Rg^* + e \rightarrow Rg + e$ $Hg + e \rightarrow Hg^* + e$ $Hg + e \rightarrow Hg^+ + 2e$ $Hg^* + e \rightarrow Hg + e$ $Rg^* + Rg^* \rightarrow Rg^+ + Rg + e$ $Rg^* + Hg \rightarrow Hg^+ + Rg + e$ $Rg^* + Hg^* \rightarrow Hg^+ + Rg + e$ $Rg^+ + Hg \rightarrow Hg^+ + Rg$ $Rg^+ + Hg^* \rightarrow Hg^+ + Rg$ $Rg^+ + Rg \rightarrow Rg + Rg^+$ $Hg^+ + Hg \rightarrow Hg + Hg^+$ $Hg^* + Hg^* \rightarrow Hg^+ + Hg + e$ $Rg^* \rightarrow Rg + hv$ $Hg^* \rightarrow Hg + hv$



MONTE CARLO RADIATION TRANSPORT MODULE

- Monte Carlo photon pseudo-particles are launched from locations proportional to Hg^{*} density.
- Trajectories are tracked accouting for absorption/emission based on Voight profile.
- Null cross section techniques account for variations in absorber and perturber densities, collision frequency and gas temperature.
- Partial frequency redistribution of emitted photons.
- Isotope shifts and fine structure splitting.
- Effective lifetimes (residence times) of photons in plasma and exit spectra are calculated.
- Rate constant of radiative reaction decreased by the trapping factor (ratio of effective to natural photon lifetime for given transition).

- Diameter 9 cm
- Height 8 cm
- Initial pressure 500 mTorr
- Initial temperature 375 K
- Power 50 W
- Frequency 10 Mhz
- Initial Ar mole fraction \cong 0.97
- Initial Hg mole fraction (ground state) ≅ 0.03



ELECTRON DENSITY AND TEMPERATURE



• The electron density goes up with trapping due to more ionization processes, while the temperature becomes more localized due to a reduced skin depth.

EFFECT OF RADIATION TRAPPING

- Radiation trapping leads to an increased lifetime for the [Hg*] atoms.
- There are more super-elastic collisions and hence most lower energy electrons gain an additional amount of energy (≅ 5 eV).



- This leads to a "bulge" in the lower energies in the EED, which is smoothed out by other inelastic collisions.
- However, since the total power deposited is a constant, higher rate of dissipation at low energies produces a decrease in the EED at higher energies.

VARIATION WITH SPATIAL POSITION



- The EED in the skin depth has a longer tail due to more stochastic heating.
- Diffusion cooling near sheath of the opposite wall depletes high energy electrons.

VARIATION WITH POWER DEPOSITED

- The EEDs shown here are for an untrapped lifetime of 125 ns $(6^{3}P_{1} \rightarrow 6^{1}S_{0}$ transition).
- As the power increases, we have more electrons in the plasma:

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=1.8 \times 10^{12} \text{ cm}^{-3} (100 W)
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- This increased electron density leads to more Maxwellian EEDs (both trapped and untrapped) for deposition of 100 W.
- The two-temperature distribution seen at 50 W is not seen at 100 W.

VARIATION WITH POWER DEPOSITED (10 ns LIFETIME)

- The EEDs shown here are for a fictitious transition having the same energy difference as the $6^{3}P_{1} \rightarrow$ $6^{1}S_{0}$ transition, but with a vacuum lifetime of 10 ns.
- The increased power deposition in the plasma leads to a doubling of electron heating in the skin depth.
- $n_e = 1.6 \times 10^{11} \text{ cm}^{-3} (50 \text{ W})$ = 2.1 x 10¹¹ cm⁻³ (100 W)



- This heating leads to a tail in the untrapped EED with increase in power.
- In the case of 10 ns radiative lifetime, we see that the electric field acceleration dominates the EED, while in the 125 ns case, e-e collisions determine the EED at high powers.

EFFECT OF ICP FREQUENCY



- With increase in frequency, the untrapped distributions do not change.
- However, the trapped EEDs are significantly different in the three cases, with lower frequencies showing a longer tail in the distribution.

EFFECT OF ICP FREQUENCY (contd.)



- There is more stochastic heating in the trapped EED at skin depth.
- At low frequencies, this heating offsets the reduction in the tail due to trapping.

SUMMARY

- A Monte Carlo Resonance Radiation Transport Model has been interfaced with a plasma equipment model to model the effects of radiation trapping on EEDs.
- Radiation trapping is seen to affect the bulk as well as the tail of the EED due to enhancement of super-elastic collisions.
- The EED becomes more Maxwellian with increase in power at high radiative lifetimes. At lower lifetimes, stochastic heating dominates the tail of the EED.
- Increased skin depth electron heating at lower frequencies for trapped distributions significantly changes the tail of the EED.